

Homework



due 2017-05-04

Exams

final: Monday, May 8, 2017, 13:00-14:59

XXII. BOSE-EINSTEIN CONDENSATION [(4+6+8) PTS]

The grand canonical potential for an ideal (non-relativistic) Bose gas is given by

$$J(T, \mu, V) = T \ln(1 - z) - T \frac{V}{\lambda_\beta^3} \text{Li}_{5/2}(z),$$

where $z = e^{\mu/T}$, $\lambda_\beta = h/\sqrt{2\pi m T}$, and $\text{Li}_s(z)$ is the polylogarithm function.

- Use the potential J and calculate the entropy $S(T, V, \mu) = -(\frac{\partial J}{\partial T})_{V, \mu}$ above and below the transition point of the Bose condensation. Show that S is continuous at the transition point and vanishes for $V = 0$ and $T = 0$.
- Calculate the specific heat

$$C_V(T, V, z) = T \left(\frac{\partial S}{\partial T} \right)_{V, N}$$

above and below the transition. In the following the volume is fixed. Notice, that the partial derivative of $S(T, V, \mu)$ needs to be calculated while keeping N constant, i.e.,

$$\left(\frac{\partial S}{\partial T} \right)_{V, N} = \left(\frac{\partial S}{\partial T} \right)_z + \frac{\partial S}{\partial z} \left(\frac{\partial z}{\partial T} \right)_N.$$

The derivative $\partial z / \partial T$ is obtained by applying the operator $(\partial / \partial T)_N$ on both sides of

$$N(T, V, \mu) = - \left(\frac{\partial J}{\partial \mu} \right)_{T, V} = \frac{z}{1 - z} + \frac{V}{\lambda_\beta^3} \text{Li}_{3/2}(z). \quad (1)$$

Is the specific heat continuous at the transition point?

- Calculate the inverse isothermal compressibility

$$K_T^{-1} = -V \left(\frac{\partial p}{\partial V} \right)_{T, N}$$

above and below the transition. Show that K_T diverges at the transition point. Notice again, that the derivative is to be calculated for constant N (T is fixed):

$$p \approx \frac{T}{\lambda_\beta^3} \text{Li}_{5/2}(z) \Rightarrow \left(\frac{\partial p}{\partial V} \right)_N = \left(\frac{\partial p}{\partial V} \right)_z + \frac{\partial p}{\partial z} \left(\frac{\partial z}{\partial V} \right)_N.$$

In order to calculate $(\frac{\partial V}{\partial z})_{T, N}$ solve eq. (1) for V . Use the above results to calculate K_T as function of the specific volume $v = V/N$ near the critical volume $v_c(T) = \lambda_\beta^3 / \text{Li}_{3/2}(1)$ at which the Bose condensation occurs, i.e., for $v - v_c \ll v_c$. Use the series expansion

$$\text{Li}_{3/2}(z) = \text{Li}_{3/2}(1) (1 - 1.36\sqrt{1 - z} + \dots).$$

If you calculated all correctly, S is continuous at the transition point and vanishes at $v = 0$ (or $T = 0$). Furthermore, $C_V(T, V, z)$ has a jump in the first derivative and K_T diverges at the transition point. Therefore, the Bose-Einstein condensation (BEC) is a phase transition of second order, if one considers the mixed phase of gas and condensate in the region $v < v_c$ as a new phase. On the other hand, if the new phase is only the phase when all particles are in the condensate at $T = 0$ (this state corresponds to the specific volume $v = 0$, since there is no repulsion), the BEC is a phase transition of first order for all isotherms.

XXIII. IDEAL FERMIONS AT LOW TEMPERATURES AND SOMMERFELD EXPANSION [(5+9+3) PTS]

Here you should apply the Sommerfeld expansion to calculate properties of Fermions at low temperatures (see lecture).

- a) Calculate $E(T, N, V)$ up to order $\mathcal{O}(T^2)$ for a given single particle density of states (DoS) $\rho(\epsilon)$. Derive from that the result for the specific heat (in order $\mathcal{O}(T)$):

$$C_V(T, N, V) = \left(\frac{\partial E}{\partial T} \right)_{V, N} = \frac{\partial E}{\partial T} + \frac{\partial E}{\partial \mu} \left(\frac{\partial \mu}{\partial T} \right)_{V, N}$$

- b) Derive an expression for the single particle DoS $\rho(\epsilon)$ in d dimensions for non-relativistic Fermions, with energy-momentum relation $\epsilon = \mathbf{p}^2/(2m)$. What is special for $d = 2$? Use this expression to calculate the energy, E , (order $\mathcal{O}(T^2)$), heat capacity, C_V , (order $\mathcal{O}(T)$), and pressure, $p(T, N, V) = -(\partial E / \partial V)_{N, T}$, (order $\mathcal{O}(T^2)$).
- c) Calculate the Fermi-energy and heat capacity for relativistic Fermions with linear energy-momentum relation, $\epsilon_{\mathbf{p}} = c|\mathbf{p}|$ (e.g., electrons in a white dwarf star).